New developments in numerical studies of quantum time-evolution based on the real-time path integral

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Ref.)

[1] JN, Katsuta Sakai, Atis Yosprakob, JHEP 09 (2023) 110, e-Print: 2307.11199 [hep-th]
[2] Chien-Yu Chou, JN, e-Print: 2407.17724 [gr-qc]
[3] JN, Hiromasa Watanabe, e-Print: 2408.16627 [quant-ph] Two approaches to quantum time-evolution

1) Schrödinger equation



advantageous for:

- applications to large numbers of d.o.f. (including QFT)
- extensions to QG (diffeomorphism) and matrix model (nonperturbative strings)
- possible applications of powerful <u>Monte Carlo simulation</u> (important sampling)

Progress in real-time path integral

 \blacktriangleright quantum mechanics : $\Psi(x_{f}, t_{f}) = \int_{x(t_{f})=x_{f}} \mathcal{D}x(t) \Psi(x(t_{i}), t_{i}) e^{iS[x(t)]/\hbar}$ ("time" is one of the dynamical variables) quantum gravity $\Psi[h] = \int \mathcal{D}g_{\mu\nu} e^{iS[g]/\hbar}$ nonperturbative formulation of string theory IKKT matrix model ("time" is an emergent concept) $Z = \int dA_{\mu} \left(e^{iS[A]} \right)$ the oscillating behavior

conceptual problem : How to define the oscillating integral

Picard-Lefschetz theory

technical problem : How to overcome the sign problem in MC sim.

 the generalized (Lefschetz) thimble method (including the "Worldvolume approach" of Fukuma-Matsumoto ('20))

Plan of the talk

- 0. Introduction
- 1. Generalized thimble method and the Picard-Lefschetz theory
- 2. A new picture of quantum tunneling
- 3. Quantum tunneling at the beginning of the universe
- 4. Quantum decoherence from saddle points
- 5. Summary and discussions

1. Generalized thimble method and the Picard-Lefschetz theory

The generalized thimble method

Alexandru, Basar, Bedaque, Ridgway, Warrington, JHEP Fukuma's talk on Sept.?? 1605 (2016) 053 $Z = \int_{\mathbb{D}^N} dx \, e^{-S(x)}$ $S(x) \in \mathbb{C}$ Lefschetz thimble $(\tau \rightarrow \infty)$ \boldsymbol{Z} $z(x;\tau)$ deformed integration contour - saddle point : $\frac{\partial S(z)}{\partial z} = 0$ anti-holomorphic gradient flow original integration contour $(x \in \mathbb{R}^N)$ \mathcal{X} Solve $\frac{\partial}{\partial \sigma} z_k(x; \sigma) = \frac{\partial S(\overline{z(x; \sigma)})}{\partial z_k}$ from $\sigma = 0$ to $\sigma = \tau$ with the initial condition $z(x; 0) = x \in \mathbb{R}^N$ One obtains a one-to-one map from x to $z(x; \tau)$, which defines a contour deformation. (Cauchy's theorem)

An important property of the gradient flow





Real part of the action increases along the flow, while the imaginary part is kept constant.



The imaginary part tends to become constant as the flow time increases. (Sign problem is solved !)



An oscillating integral can be made well defined uniquely. No ambiguity in the choice of integration contour.

2. A new picture of quantum tunneling

JN, Katsuta Sakai, Atis Yosprakob,

"A new picture of quantum tunneling in the real-time path integral from Lefschetz thimble calculations" JHEP 09 (2023) 110, e-Print: 2307.11199 [hep-th]

Time-evolution of the wave function

$$\Psi(x_{\mathsf{f}}, t_{\mathsf{f}}) = \int_{x(t_{\mathsf{f}}) = x_{\mathsf{f}}} \mathcal{D}x(t) \Psi(x(t_{\mathsf{i}}), t_{\mathsf{i}}) e^{iS[x(t)]}$$

$$S[x(t)] = \int dt \left\{ \frac{1}{2} m \left(\frac{dx}{dt} \right)^2 - V(x) \right\}$$
$$V(x) = \alpha (x^2 - 1)^2 \qquad \alpha = 2.5$$

$$\Psi(x, t_{i}) = \exp\left\{-\frac{1}{4\sigma^{2}}(x-b)^{2}\right\}$$
$$\sigma = 0.3, \quad b = -1$$
$$x_{f} = 1$$
$$T \equiv t_{f} - t_{i} = 2$$

Discretize the time as:

$$x_n = x(t_n)$$

$$t_n = \frac{n}{N}T \qquad (n = 0, \dots, N)$$

$$N = 20$$



Results of GTM (with flow time integration *a la* Fukuma-Matsumoto)



Introducing momentum in the initial state



A new understanding of quantum tunneling $\Psi(x_{f}, t_{f}) = \int \mathcal{D}x(t) \Psi(x(t_{i}), t_{i}) e^{iS[x(t)]/\hbar}$

 $x(t_{\rm f}) = x_{\rm f}$ initial wave function $\Psi(x, t_{\rm i}) = \varphi(x) e^{ipx/\hbar}$

 $\varphi(x)$ is assumed to have a finite support $\Delta \equiv [x_{\min}, x_{\max}]$

 $\hbar \to 0$ Classical EOM $\frac{\delta S[x(t)]}{\delta x(t)} = 0$

If real x(t) exists,
 it is a relevant & dominant saddle.

 If real x(t) does not exist, the relevant saddle with min. ImS(> 0) dominates.

|prob. amplitude| $\sim e^{-\mathrm{Im}S[x^{\star}]/\hbar}$ (instanton-like suppression) Boundary condition $x(t_i) \in \Delta, \quad \dot{x}(t_i) = \frac{p}{m}$ $x(t_f) = x_f$

real trajectory emergence of classical motion

<mark>complex trajectory</mark>

semi-classical description of quantum tunneling

Can be observed by using the weak measurement

3. Quantum tunneling at the beginning of the universe

Chien-Yu Chou, JN, "Monte Carlo studies of quantum cosmology by the generalized Lefschetz thimble method" e-Print: 2407.17724 [gr-qc]

How did the Universe begin ?



A. Vilenkin

"Creation of Universes from Nothing" Phys.Lett.B 117 (1982) 25



tunneling from nothing



J. B. Hartle and S. W. Hawking "Wave function of the Universe" Phys. Rev. D 28,(1983) 2960



no-boundary wavefunction

Quantum Cosmology

Issues in quantum cosmology

 Vilenkin's proposal seems to have instability in fluctuations due to the "wrong" Wick rotation.



• Hartle-Hawking's proposal seems to be **incompatible** with the inflation scenario since it favors $\Lambda = 0$.

Which is the relevant saddle point ?

mini-superspace model

Halliwell-Louko, Phys.Rev.D 39 (1989) 2206

Assuming homogeneous, isotropic, closed space-time

 $ds^{2} = a^{2}(\eta)(-N(\eta)^{2}d\eta^{2} + d\Omega_{3}^{2})$ η : conformal time scale factor lapse function

Einstein-Hilbert action

Instein-Hilbert action

$$S_{\mathsf{EH}}[a,N] = 6\pi^2 \int d\eta \left\{ -\frac{1}{N} \left(\frac{da}{d\eta} \right)^2 + NV(a) \right\} \qquad V(a) = a^2 - \frac{\Lambda}{3} a^4$$

change of variables: $\begin{cases} q = a^2 \\ dn = a^{-2}(t)dt \end{cases}$

$$ds^{2} = -\frac{N^{2}}{q(t)}dt^{2} + q(t) d\Omega_{3}^{2} \qquad N = \text{const. (reparam. inv.)}$$
$$S_{\mathsf{EH}}[q, N] = 6\pi^{2} \int_{0}^{1} dt \left\{ -\frac{1}{4N} \left(\frac{dq}{dt}\right)^{2} + N \left(1 - \frac{\Lambda}{3}q\right) \right\}$$

q(t) can be integrated out by the Gaussian integral Integration over N has ambiguity in the choice of contour.

Picard-Lefschetz theory \rightarrow Vilenkin's saddle becomes relevant.

Feldbrugge-Lehners-Turok, Phys.Rev.D 95 (2017) 10, 103508, 1703.02076 [hep-th]

Simulating quantum cosmology



Next step : Add tensor modes and investigate the instability issue.

4. Quantum decoherence from saddle points

JN, Hiromasa Watanabe, "Quantum decoherence from saddle points" e-Print: 2408.16627 [quant-ph]

Couple the system to an environment

$$L = L_{\mathcal{S}} + L_{\mathcal{E}} + L_{\text{int}}$$

Caldeira-Leggett ('83)

$$L_{\mathcal{S}} = \frac{1}{2} M \dot{x}(t)^2 - \frac{1}{2} M \omega_b^2 x(t)^2 ,$$

$$L_{\mathcal{E}} = \sum_{k=1}^{N_{\mathcal{E}}} \left\{ \frac{1}{2} m \, \dot{q}^k(t)^2 - \frac{1}{2} m \, \omega_k^2 \, q^k(t)^2 \right\} ,$$

$$L_{\text{int}} = c \, x(t) \sum_{k=1}^{N_{\mathcal{E}}} q^k(t) ,$$

reduced density matrix after tracing out the environment



Exact results from saddle points

JN, Hiromasa Watanabe, e-Print: 2408.16627 [quant-ph]

Introducing
$$X_{\mu} = \{x_i, \tilde{x}_i, q_i^k, \tilde{q}_i^k, (\tilde{q}_0^k)_j\}$$

$$S_{\text{eff}}(x, \tilde{x}, q, \tilde{q}, \tilde{q}_0) = \frac{1}{2} X_{\mu} \mathcal{M}_{\mu\nu} X_{\nu} - C_{\mu} X_{\mu} + B$$

saddle point:
$$\bar{X}_{\mu} = \left(\mathcal{M}^{-1}\right)_{\mu\nu} C_{\nu}$$

 $X_{\mu} = \bar{X}_{\mu} + Y_{\mu}$

Integrating Y_{μ} ,

$$\rho_{\mathcal{S}}(x_{\mathsf{F}}, \tilde{x}_{\mathsf{F}}; t_{\mathsf{F}}) = \frac{1}{\sqrt{\det \mathcal{M}}} e^{-\mathcal{A}} ,$$
$$\mathcal{A} = B - \frac{1}{2} C_{\mu} \left(\mathcal{M}^{-1} \right)_{\mu\nu} C_{\nu}$$

Quantum decoherence is captured by complex saddle points. (analogous to what we have found for quantum tunneling)

Disappearance of interference pattern

 γ : coupling with environment β : inverse temperature

JN, Hiromasa Watanabe, in preparation

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 $- \gamma = 0.05$

 $--\gamma = 0.5$

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 $(N_{\mathcal{E}} = 64, \, \omega_{\text{cut}} = 2.0, \, \omega_{\text{r}} = 0, \, \sigma = 0.1)$ $ho_{\mathcal{S}}(x,x;t)$ t = 0.3t = 0.5t = 0.7t = 00.80.80.80.8t = 0.3t = 0.5t = 0.1 $\beta = 0.05 \bullet \bullet \cdot \beta = 0.1$ 0.60.60.6 $\beta = 0.5 \quad \bullet \bullet \cdot \gamma = 0$ 0.60.40.40.4 $\beta = 0.5, 0.1, 0.05$ 0.4v = -1t = 0.70.20.20.20.2 $(\gamma = 0.1)$ 0.00.0-2-20 2 -22 2 $^{-2}$ -44 0 $\mathbf{4}$ 0 4 xxxx0.80.80.80.8t = 0.3t = 0.1t = 0.50.60.60.6 0.6= 0.1 $\gamma = 0.05, 0.1, 0.5$ 0.4 0.40.40.4t = 0.7 $(\beta = 0.05)$ 0.20.20.20.20.00.0 0.0 0.0-20 $\mathbf{2}$ 4 -20 24 -20 24 -20 -4-4-4xxxx

Master eq. (with Born and Markov approximations)

$$\frac{d}{dt}\rho_{\mathcal{S}}(x,\tilde{x};t) = K(x,\tilde{x})\rho_{\mathcal{S}}(x,\tilde{x},t) , \qquad (\omega_{\mathsf{r}} \ll \omega_{\mathsf{cut}} \ll T = \beta^{-1})$$

$$K(x,\tilde{x}) = \frac{\mathsf{i}}{2M} \left(\frac{\partial^2}{\partial x^2} - \frac{\partial^2}{\partial \tilde{x}^2} \right) - \frac{\mathsf{i}}{2} M \omega_{\mathsf{r}}^2 (x^2 - \tilde{x}^2) - \gamma(x - \tilde{x}) \left(\frac{\partial}{\partial x} - \frac{\partial}{\partial \tilde{x}} \right) \left(\frac{2M\gamma}{\beta} (x - \tilde{x})^2 \right)$$
The effect of decoherence is $\propto \frac{\gamma}{\beta}$

Quantum decoherence is captured by complex saddles in the real-time path integral.

5. Summary and discussions

Summary

Quantum time-evolution includes many interesting physics.

- quantum tunneling
- beginning of the Universe
- quantum decoherence

• Real-time path integral : very useful in studying these things.

- Oscillating integral can be dealt with by the Picard-Lefschetz theory.
- These phenomena can be captured by relevant (complex) saddle points.
- Monte Carlo simulation is possible by using the generalized thimble method.

• Various applications are waiting for us!

- Quantum-to-classical transition (How does the classical world emerge?)
- Measurement problems (Schrödinger's cat)
- Instability problem of the Vilenkin's saddle (smooth beginning of the Universe)
- Matrix model (emergence of (3+1)-expanding Universe from superstring theory)
- Quantum chaos (and its relation to blackholes through holography)
- Quantum information (and its relation to AdS/CFT) etc.