QCD thermodynamics at non-zero isospin and baryon density

Bastian Brandt

Universität Bielefeld

Gergely Endrődi & Gergely Markó





Faculty of Physics



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QCD systems at non-zero chemical potential



Strongly interacting matter under extreme conditions:

- \blacktriangleright a number of systems dominated by μ_B
 - Heavy-Ion collisions
 - Neutron stars

important contribution from $\mu_Q \neq 0$

- \blacktriangleright others possibly dominated by μ_Q
 - Early Universe at large individual l_ℓ

Important for description:

need the EoS at $\mu_B,\,\mu_Q,\,\mu_S\neq 0$







Bastian Brandt



Typical method: integral method with Taylor expansion around $\mu = 0$

$$p(T,\vec{\mu}) = p(T,\vec{0}) + \sum_{n_{1,2,3}=1}^{\infty} \left. \frac{\partial^{n_1+n_2+n_3}[p(T,\vec{\mu})]}{\partial \mu_1^{n_1} \partial \mu_2^{n_2} \partial \mu_3^{n_3}} \right|_{\vec{\mu}=\vec{0}} \prod_i \frac{1}{n_i!} (\mu_i)^{n_i}$$





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"Physical" chemical potential basis:





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"Physical" chemical potential basis:

$$\mu_u = \frac{\mu_B}{3} + \frac{2\mu_Q}{3} \qquad \mu_d = \frac{\mu_B}{3} - \frac{\mu_Q}{3} \qquad \mu_s = \frac{\mu_B}{3} - \frac{\mu_Q}{3} - \mu_S$$

Limitations:

- finite radius of convergence (bounds difficult to compute)
- cannot extrapolate accross phase transitions





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Better for systems with large μ_Q : expand directly around non-zero μ_I



QCD at non-zero isospin chemical potential





[Brandt, Endrődi, Schmalzbauer '18]

EoS from canonical @ T = 0 [Detmold *et al* '12, '23, '24]



Equation of state at non-zero μ_I

II Taylor expansion to non-zero $\mu_{L,s}$

III An application: EoS at non-zero charge chemical potential



I Equation of state at non-zero μ_I

Equation of state at $\mu_I \neq 0$







Equation of state at $\mu_I \neq 0$



$$p(T, \mu_I) = p(T, 0) + \Delta p(T, \mu_I)$$

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$$\blacktriangleright \Delta I(T,\mu_I) = \int_0^{\mu_I} d\mu'_I \Big(T\frac{\partial}{\partial T} - 4\Big) n_I(T,\mu'_I) + \mu_I n_I(T,\mu_I)$$

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use all spline interpolations which provide a "good" description of $n_{I} \,$



Equation of state at $\mu_I \neq 0$



$$p(T, \mu_I) = p(T, 0) + \Delta p(T, \mu_I) \qquad I(T, \mu_I) = I(T, 0) + \Delta I(T, \mu_I)$$
[Borsanyi *et al* '13, Bazavov *et al* '14] to be determined

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need to compute:

$$n_{I} = \frac{T}{V} \mathsf{Tr} \Big[M^{-1} \frac{\partial M}{\partial \mu_{I}} \Big] \equiv \frac{T}{V} c_{I}$$

M: fermion matrix





$$D = \gamma_{\mu} D_{\mu} + m_{ud} + \gamma_{0} \tau_{3} \mu_{I}$$

$$SU_{V}(2) \xrightarrow{\bullet} U_{Q}(1) \xrightarrow{\bullet} \varnothing$$

$$explicit \qquad \text{spontaneous}$$

$$\mu_{I} \neq 0 \qquad \mu_{I} \geq m_{\pi}/2$$

- cannot observe spontaneous symmetry breaking in finite ${\cal V}$
- low mode in simulations







 $SU_V(2) \xrightarrow{\bullet} U_Q(1) \xrightarrow{\bullet} \varnothing$ explicit $\mu_I \neq 0$ pionic source λ

 need to break symmetry explicitly
 → introduce regulator: ~ λ pionic source [Kogut, Sinclair '02]
 physical results: extrapolate λ → 0 main task for analysis

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$$D = \gamma_{\mu} D_{\mu} + m_{ud} + \gamma_0 \tau_3 \,\mu_I + i\gamma_5 \tau_2 \,\lambda$$

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facilitated by improvement program:
 [Brandt, Endrődi, Schmalzbauer '18]

- leading order reweighting
- valence quark improvement for light quarks

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Up to now: consider 1st derivatives - densities

$$O \equiv \mathrm{Tr}\Big[M^{-1}\hat{O}\Big] \qquad c_X: \ \hat{O} = \frac{\partial M}{\partial \mu_X} \quad \text{with} \quad M = D^{\dagger}(\mu)D(\mu) + \lambda^2$$

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CRC-TR

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Results for the EoS at $\mu_I \neq 0$ – $N_t = 8$







II Taylor expansion to non-zero $\mu_{L,s}$

Taylor expansion around $\mu_I \neq 0$



Extension to μ_L , $\mu_s \neq 0$:

$$p(T,\vec{\mu}) = p(T,\mu_I,0,0) + \sum_{n,m=1}^{\infty} \frac{1}{n!\,m!} \left. \frac{\partial^n \partial^m [p(T,\vec{\mu})]}{\partial \mu_L^n \,\partial \mu_s^m} \right|_{\mu_L,\mu_s=0} \left(\mu_L \right)^n \left(\mu_s \right)^m$$



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Leading order: only non-zero coefficients are

$$\chi_2^L(T,\mu_I) \equiv \left. \frac{\partial^2[p(T,\vec{\mu})]}{\partial \mu_L^2} \right|_{\mu_L,\mu_s=0} \& \quad \chi_2^s(T,\mu_I) \equiv \left. \frac{\partial^2[p(T,\vec{\mu})]}{\partial \mu_s^2} \right|_{\mu_L,\mu_s=0}$$

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• Generically:
$$\chi_2^X = \frac{T}{V} \Big[\underbrace{\langle c_{XX} \rangle}_{\text{connected}} + \underbrace{\langle (c_X)^2 \rangle - \langle c_X \rangle^2}_{\text{disconnected}} \Big]$$

with
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 M : fermion matrix
 $c_{XX} = \frac{\partial c_X}{\partial \mu_X} = \operatorname{Tr}\left[M^{-1}\frac{\partial^2 M}{(\partial \mu_X)^2}\right] + \operatorname{Tr}\left[M^{-1}\frac{\partial M}{\partial \mu_X}M^{-1}\frac{\partial M}{\partial \mu_X}\right]$



$$\chi_2^X = \frac{T}{V} \left[\left\langle c_{XX} \right\rangle + \left\langle (c_X)^2 \right\rangle - \left\langle c_X \right\rangle^2 \right]$$

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$$c_{XX} = \mathsf{Tr}\Big[M^{-1}\frac{\partial^2 M}{(\partial\mu_X)^2}\Big] + \mathsf{Tr}\Big[M^{-1}\frac{\partial M}{\partial\mu_X}M^{-1}\frac{\partial M}{\partial\mu_X}\Big]$$

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$$\begin{split} \begin{array}{lcl} C_{12} & \equiv & \mathrm{Tr}\Big[M^{-1}\hat{O}^{(1)}M^{-1}\hat{O}^{(2)}\Big] \\ & \approx & \left\langle \sum_{n,m=0}^{N}\frac{\mathcal{O}_{nm}^{(1)}}{\xi_{m}^{2}+\lambda^{2}}\frac{\mathcal{O}_{nm}^{(2)}}{\xi_{n}^{2}+\lambda^{2}}\right\rangle \end{split}$$



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treat as above

0

improvement term:

$$\delta_{C_{12}}^{N} = \sum_{n,m=1}^{N} \mathcal{O}_{nm}^{(1)} \mathcal{O}_{mn}^{(2)} \left(\frac{1}{\xi_{n}^{2} + \lambda^{2}} \frac{1}{\xi_{m}^{2} + \lambda^{2}} - \frac{1}{\xi_{n}^{2}} \frac{1}{\xi_{m}^{2}} \right)$$



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10 ----

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λ -extrapolated Taylor coefficents @ $N_t=8$



• results for $\chi_2^L(T, \mu_I)$ using standard λ -extrapolations:



Large uncertainties for $\chi_2^L(T, \mu_I)$ in the BEC phase!

 $\blacktriangleright \chi_2^s(T,\mu_I)$ not affected (no source parameter) 0.6T = 123 MeVT = 147 MeV0.4 0.4 χ^s_2/m^2_{π} c_2^s/m_{π}^2 0.20.20 ίō 0.20.40.6 0.8 ίō 0.20.4 0.6 0.8 μ_I/m_{π} μ_I/m_{π}

Improved computation method



Observation: equal connected parts in μ_L and μ_I derivatives

 $c_{LL} = c_{II}$

Can we use that somehow?



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Can we use that somehow?

for EoS computation: model independent spline interpolation of n_I [Brandt, Cuteri, Endrődi '22]

 \rightarrow know μ_I dependence of n_I

compute $\frac{\partial n_I}{\partial \mu_I} = \chi^I_2(T,\mu_I)$ analytically from spline interpolation



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• finally compute $\chi_2^L(T,\mu_I)$ using

$$\chi_2^L(T,\mu_I) = \underbrace{\chi_2^I(T,\mu_I)}_{\lambda=0} + \frac{T}{V} \Big[\underbrace{\langle (c_L)^2 \rangle - \langle c_L \rangle^2 - \{\langle (c_I)^2 \rangle - \langle c_I \rangle^2 \}}_{\lambda \neq 0} \Big]$$

(λ -extrapolation for disconnected typically better behaved)



• comparison of the methods for $\chi_2^L(T,\mu_I)$ in BEC:



 \rightarrow reduced uncertainties for $\chi_2^L(T,\mu_I)$ in the BEC phase!

Results for Taylor coefficients @ $N_t = 8$



To extend the EoS: interpolate Taylor expansion coefficients as for EoS: [Brandt, Cuteri, Endrődi '22] [Brandt, Endrődi '16] use spline interpolations with Monte-Carlo generated nodepoints

• constraint: $\chi_2^L(T,\mu_I) \& \chi_2^s(T,\mu_I)$ vanish at T=0



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III An application: EoS at non-zero charge chemical potential

EoS for early Universe @ large l_ℓ



early Universe: sum of lepton flavour asymmetries is constrained $|l_e + l_\mu + l_\tau| < 0.012$ [Oldengott, Schwarz '17]

but: for large individual l_ℓ

 \longrightarrow large μ_Q and small μ_B & μ_S along trajectories



• model equation of state – matched to lattice at pure μ_I

[Vovchenko, et al '20]

▶ Taylor expansion from $\vec{\mu} = 0$ − no pion condensation

[Middeldorf-Wygas, et al '20]



can provide a full lattice EoS

non-zero μ_Q in the isospin basis:



assumption: BEC phase boundary does not change drastically

Conclusions



- Simulations at µ_I ≠ 0: offer a novel expansion point to explore (µ_B, µ_Q, µ_S) space
- λ -extrapolations necessary:
 - $\longrightarrow \text{ large uncertainties for } \chi_2^L(T,\mu_I) \\ \text{ in BEC phase }$
- alternative: computation via χ^I₂(T, μ_I) obtained from spline interpolation of n_I
 - improves uncertainties

some details still to be understod

 application and outlook: compute EoS at non-zero μ_Q and in its vicinity

